

General Disclaimer

One or more of the Following Statements may affect this Document

- This document has been reproduced from the best copy furnished by the organizational source. It is being released in the interest of making available as much information as possible.
- This document may contain data, which exceeds the sheet parameters. It was furnished in this condition by the organizational source and is the best copy available.
- This document may contain tone-on-tone or color graphs, charts and/or pictures, which have been reproduced in black and white.
- This document is paginated as submitted by the original source.
- Portions of this document are not fully legible due to the historical nature of some of the material. However, it is the best reproduction available from the original submission.

X-551-71-122

PREPRINT

NASA TM X- 65533

**CLOSED FORM SATELLITE
TRACKING DATA CORRECTIONS
FOR AN ARBITRARY
TROPOSPHERIC PROFILE**

JOHN W. MARINI

MARCH 1971



GODDARD SPACE FLIGHT CENTER

GREENBELT, MARYLAND

FACILITY FORM 602

N71-25689 (ACCESSION NUMBER)	(THRU)
54 (PAGES)	Q3 (CODE)
TMX-65533 (NASA CR OR TMX OR AD NUMBER)	30 (CATEGORY)



X-551-71-122

CLOSED FORM SATELLITE TRACKING DATA CORRECTIONS
FOR AN ARBITRARY TROPOSPHERIC PROFILE

John W. Marini

March 1971

Goddard Space Flight Center

Greenbelt, Maryland

CLOSED FORM SATELLITE TRACKING DATA CORRECTIONS
FOR AN ARBITRARY TROPOSPHERIC PROFILE

John W. Marini

ABSTRACT

The formulas commonly used to correct satellite tracking data for the effects of atmospheric refraction are based, for the sake of mathematical convenience, on simple model atmospheric profiles that are not physically realistic. A method is given here for deriving correction formulas that do not suffer from this limitation, and can be used in precision orbit calculations. The refractivity of the atmosphere is assumed to have spherical symmetry, but may have any given vertical profile. The method was tested for numerical accuracy by application to the simple exponential profile, and the corrections calculated agreed closely at all elevation angles from 0 to 90 degrees with those obtained by double-precision ray-tracing. The error was in all cases less than 1%, and less than 1/3% above 1 degree elevation.

The method has been used to obtain improved correction formulas, and these will be published in a separate report.

Page intentionally left blank

CONTENTS

	Page
1. INTRODUCTION	1
1.1 Background	1
1.2 Related Work	2
1.3 Satellite Tracking Data and Correction Formula	2
2. GEOMETRY AND NOTATION	4
3. INPUTS TO CORRECTION FORMULAS	8
4. CORRECTIONS USING KNOWN ANGLE OF ARRIVAL	10
4.1 Formula for Elevation Error	10
4.2 A Digression on the Approximations Used	11
4.3 Expansions of the Bending Integral	11
4.4 Form of the Approximation	14
4.5 Formula for Range Error	15
5. CORRECTIONS USING KNOWN ELEVATION ANGLE	18
5.1 Elevation Correction	18
5.2 Range Correction	21
6. EXAMPLE USING AN EXPONENTIAL PROFILE	22
6.1 Known Arrival-Angle	22
6.2 Known Elevation-Angle	23
6.3 Numerical Examples and Comparisons	24
7. SUMMARY	28

CONTENTS--(continued)

	Page
8. ACKNOWLEDGEMENT	29
REFERENCES	29
Appendix 1. Expression for ϕ in Terms of τ	32
Appendix 2. Asymptotic Expansion of I, J, and K	34
Appendix 3. Integrals for I, J, and K, and Expansions for Small Values of α	36
Appendix 4. Approximation for the Geometrical Difference	39
Appendix 5. Range-Rate Correction	41
Appendix 6. Calculation of $U'(\theta)$	42
Appendix 7. Correction Equations for an Exponential Profile, Arrival Angle Known	44
Appendix 8. Correction Equations for an Exponential Profile, Elevation Angle Known	47
Appendix 9. Improved Fit	50

CORRECTING SATELLITE TRACKING DATA FOR A SPHERICALLY-SYMMETRIC ATMOSPHERE

1. INTRODUCTION

1.1 Background

In the radio tracking of a satellite by a ground station, measurements are made of the satellite elevation, range, and range-rate. The elevation can be determined by a measurement of the angle-of-arrival of a radio wave from the satellite.

The range is obtained from a measurement of the delay of signals propagated between the ground station and the satellite, while the range-rate is obtained by counting cycles of the Doppler-shifted received signals over a prescribed period of time, or, alternatively, by noting the time interval required to count a prescribed number of cycles.

The passage of the radio waves through the atmosphere of the earth introduces into these measurements errors that may require correction. Both tropospheric and ionospheric effects are present. However, if the radio frequency used is sufficiently high, the ionospheric correction may either be neglected or superposed on the tropospheric correction [1].

1.2 Related Work

A formula for the bending of a radio ray in an exponential atmosphere as a function of the angle-of-arrival has been given by Thayer [2]. Freeman [3] has

obtained, for an exponential atmosphere, a range-error correction as a function of the elevation angle of the satellite. The correction is of the first order in the surface refractivity. Reichley [4] has obtained both second order elevation-error and range-error correction formulas as functions of the elevation angle. His formulas can be applied to profiles other than exponential. Hopfield [5], by neglecting path curvature, has obtained range and range-rate corrections for the two-quartic tropospheric refractivity profile. Rowlandson and Moldt [6] have derived closed-form corrections for an exponential atmosphere.

1.3 Satellite Tracking Data and Correction Formulas

The satellites tracked are ordinarily above the region, which extends from the ground to about 70 km in altitude, where most of the radio-ray bending caused by the nonionized atmosphere of the earth takes place.

In a typical pass of satellite over a ground station, the satellite might be under observation for a number of minutes. During this period of time, angle-of-arrival range and range-rate measurements are taken periodically at a rate, perhaps, of one set of data per second [7]. Consequently, a large volume of data is generated, which must be processed automatically by electronic computer.

To be usable in a practical sense, the equations employed to correct this data for the effects of the atmosphere should require a minimum of computer time.

It is not easy, however, to derive correction equations that, on the one hand, accurately represent the atmospheric model assumed and, on the other hand, are suitable for use in the processing of large quantities of data. The need to obtain a mathematically tractable formula was a consideration in the choice of the fourth (quartic) power in the two-quartic profile [5]. The same mathematical difficulty has also resulted in the existence and use of a multitude of approximate correction formulas [8]. Most of these are in agreement at high elevation angles, but disagree at low angles, apparently because of differing approximations used, even when the same profile was initially assumed. Fortunately this discrepancy is not too serious since most satellite tracking measurements are taken at the higher elevation angles. Nevertheless, it is possible to obtain, for any assumed profile, corrections that are both accurate and yet suitable for use on tracking data. To do so it is necessary to take advantage of the circumstance that the corrections ordinarily are exercised repeatedly using varying values of elevation angle and range but with a fixed atmospheric profile. Consequently, initial or "pre-pass" calculations which involve only the atmospheric conditions at the time of the satellite pass and which are independent of satellite position may be lengthy without causing a significant percentage increase in the total computer time per satellite pass. For efficiency in calculation, therefore, the mathematical formulation of the corrections should be such that quantities functionally dependent on the

parameters of the atmospheric profile are separated from those dependent on elevation and range.

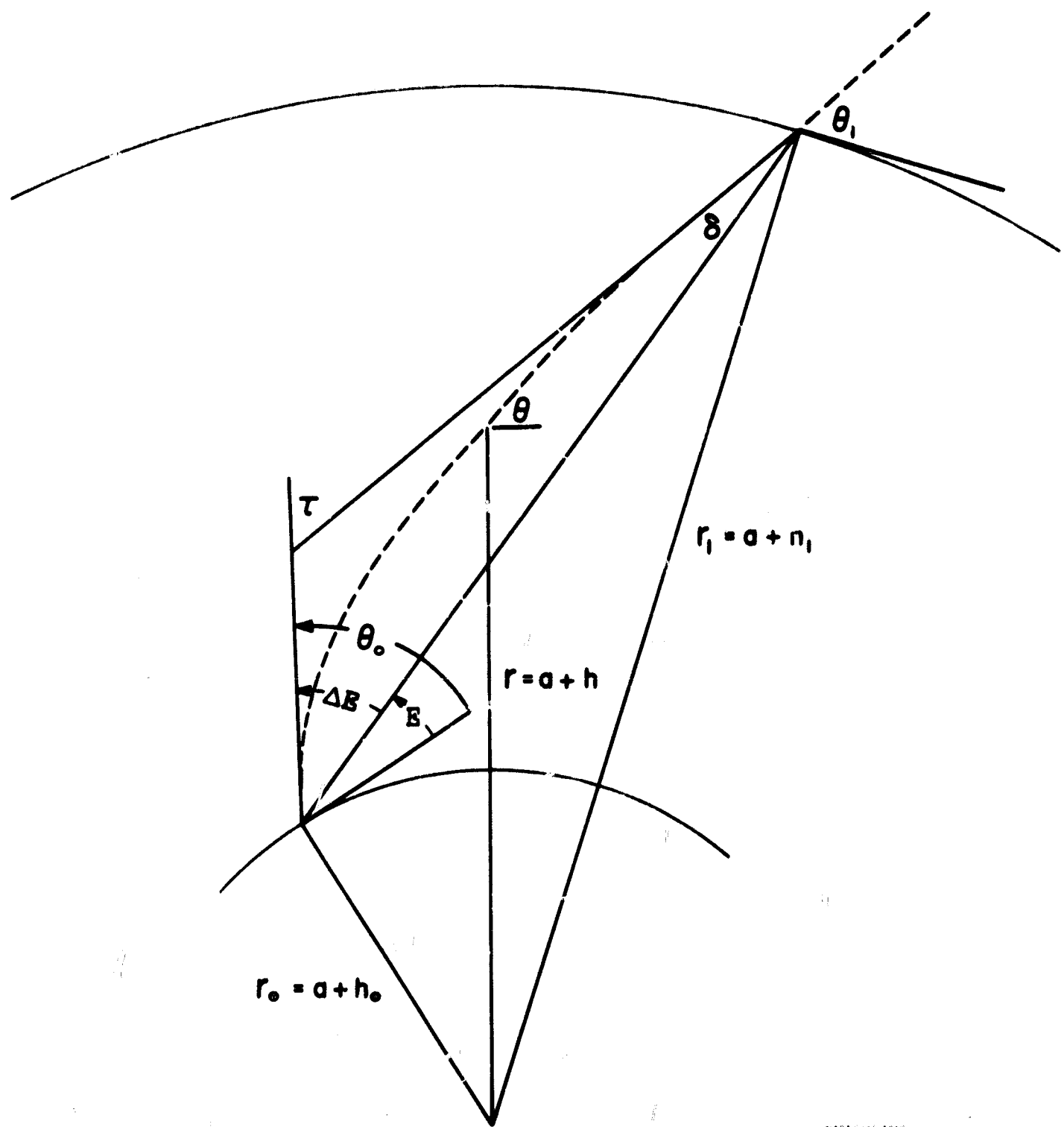
Such a separation can be effected by expansions in rational functions of the sine of the elevation angle and negative powers of the range. The coefficients in these expansions will depend on the atmospheric profile alone, and can be calculated in advance of the satellite pass.

2. GEOMETRY AND NOTATION

The geometry involved is shown in Figure 1. The earth is taken to be spherical with a radius 'a' nominally equal to 6369.95 kilometers. The tracking station is located at a distance r_0 from the center of the earth, and at a height h_0 above sea level. The satellite is at a distance r_1 from the center of the earth, and at a height h_1 . The radio ray path between the satellite and the station is shown as a dotted line. The distance between the center of the earth and a given point on the ray path is 'r'. The height of the point above sea level is 'h', and the elevation of the ray path at the point is θ .

The angle-of-arrival is the angle θ_0 above the local horizontal at the ground station. The angle ΔE is the elevation error, or difference between the angle-of-arrival and the true elevation angle E of the satellite with respect to the ground station.

$$\Delta E = \theta_0 - E \quad (1)$$



NASA/CPL TRGS
MISSION & TRAJECTORY ANALYSIS DIVISION
BRANCH MSAB DATE 3 9 71
BY M... PLOT NO. 1263

Figure 1. Satellite and Tracking Station Geometry

and τ is the total bending of the radio path. The refractive index n is assumed to depend only on the height h above the surface of the earth. The radio refractivity

$$N(h) = 10^6 [n(h) - 1] \quad (2)$$

will be used in normalized form, and a normalized height above the tracking station will be employed. Taking the refractive index and the refractivity at the tracking station to be

$$n_0 = n(h_0) \quad (3)$$

$$N_0 = N(h_0) \quad (4)$$

respectively, and taking the effective height of the troposphere above the tracking station to be

$$H = (1/N_0) \int_{h_0}^{\infty} N(h) dh \quad (5)$$

the normalized height above the tracking station

$$x = (h - h_0) / H \quad (6)$$

is defined. In terms of this variable, the refractivity may be written as

$$N(h) = N_0 f(x) \quad (7)$$

where the normalized profile $f(x)$, which will be abbreviated as f , is

$$f = N (h_0 + H x) / N_0 \quad (8)$$

which is equal to unity at $x = 0$

$$f(0) = 1 \quad (9)$$

and integrates to unity

$$\int_0^{\infty} f \, dx = 1 \quad (10)$$

from the tracking station upward. In the particular case of an exponential profile, f is equal to $\exp(-x)$.

The radio range, or electrical distance along the ray path is designated as R_e ,

$$R_e = \int_{r_0}^{r_1} (n / \sin \theta) \, dr \quad (11)$$

the geometrical distance along the ray path is

$$R_g = \int_{r_0}^{r_1} (1 / \sin \theta) \, dr \quad (12)$$

and the straight-line distance or slant range is R . The range error is then the difference

$$\Delta R = R_e - R \quad (13)$$

The parameters

$$p = \sqrt{2H/r_0} \quad (14)$$

$$q = 10^{-6} N_0 r_0 / H = 2 \times 10^{-6} N_0 / p^2 \quad (15)$$

$$Q = q \cos^2 \theta_0 \quad (16)$$

the normalized sine of the angle of arrival

$$\alpha = (1/p) \sin \theta \quad (17)$$

the normalized sine of the elevation angle,

$$\beta = (1/p) \sin E \quad (18)$$

and the normalized inverse range

$$\rho = p r_0 / R \quad (19)$$

will be used. Typical values are $p = 0.05$ and $q = 0.25$. The value of α range from 0 on the horizon to about 20 for a wave arriving vertically downward.

3. INPUTS TO THE CORRECTION FORMULAS

The desired form of the correction formulas for the elevation error ΔE and the range error ΔR depend on the intended use. In the first case to be considered,

which applies, for example, to a tracking radar, it is assumed that measured values of the angle-of-arrival θ_0 and the radio-range R_e are available. The quantities ΔE and ΔR should then be given as functions of the variables θ_0 and R_e .

In the second case, the situation is somewhat different. It is assumed that the satellite ephemeris is known from previous tracking and orbit determination. In this case the true slant range R and the true elevation angle E will be known quite accurately as functions of time, and one or both of the corrections ΔE and ΔR may be needed to provide accurate predictions of the angle-of-arrival and/or radio-range and/or range rate either for acquisition or for comparison with values to be measured. The latter comparisons are used to improve the satellite ephemeris by an iterative process that minimizes some weighted function of the differences observed. Here ΔE and ΔR need to be expressed as functions of E and R .

In both cases above, the percentage difference between R and R_e is small. R_e is at worst about 200 meters larger than R , while R has been taken to be at least 70 km. Consequently the distinction between these quantities may be neglected in their use as inputs to the error formulas. The distinction between E and θ_0 , however, must be retained at lower elevation angles.

4. CORRECTIONS USING KNOWN ANGLE OF ARRIVAL

4.1 Formula for Elevation Error

A procedure for determining the elevation error has been given by Bean and Thayer [9]. The bending τ is first calculated by integration, and the angle δ is next calculated from the geometry of Figure 1 using the known value of τ . The elevation error is then given by

$$\Delta E = \tau - \delta \quad (20)$$

Using the notation of Section 2, the usual expression for the bending is [2]

$$\tau = 10^{-6} N_0 \cos \theta_0 I(\alpha) / p \quad (21)$$

where the bending integral is defined as

$$I(\alpha) = \int_0^{\infty} \frac{-f'}{\sqrt{x + \alpha^2 - Q(1-f)}} dx \quad (22)$$

in which it has been assumed that the satellite height h_1 is great enough to permit the upper limit to be extended to infinity.

The equation giving δ in terms of τ is derived in Appendix 1. Using (21) and (1-4) in (20), the formula for the elevation error is

$$\Delta E = 10^{-3} N_0 (1/p) \cos \theta_0 [I(\alpha) - \rho L(\alpha)] \quad \text{mrad} \quad (24)$$

with

$$L(\alpha) = 1 - \alpha I(\alpha) + \frac{1}{4} q I^2(\alpha) \quad (25)$$

4.2 A Digression on the Approximations Used

The quantities 10^{-6} $N_0 \sim 300 \times 10^{-6}$ and $p^2 \sim 0.0025$ are neglected compared to

1. However the quantity $q \sim 0.25$ cannot be so neglected, even though it contains the surface refractivity as a factor.

The quantity Q that appears in the radical in Equation (22) is not independent of satellite position because of its dependence on θ_0 . This would lead to complication of the formulas to be derived. Fortunately Q may be replaced by q with negligible error, since the term neglected thereby is small compared to the square of α

$$q \sin^2 \theta_0 (1 - f) \ll \alpha^2 = (1 - p^2) \sin^2 \theta_0$$

It will also be found that the same approximation can¹, and in some cases should², be made elsewhere in the formulas to be derived.

4.3 Expansions of the Bending Integral

In order to use the formula (24) for the elevation error a rapid method for calculating the bending integral, Equation (22), is needed. Since α is as large as 20 at high elevation angles, it is natural to expand (22) asymptotically³ in powers

1. In the asymptotic expansions that follows, when $Q = q - q p^2 \alpha^2$ is set equal to q , the term neglected is small compared to the preceding term in the series.

2. If the approximation is used in calculating $I(\alpha)$ then it should be used in calculating polynomials involving $I(\alpha)$ (for example, in eq 25) in order to obtain the correct asymptotic expansion of the polynomial.

3. The expansion is asymptotic if $q = 0$ and $f = \exp(-x)$. The nature of the expansion was not investigated for other cases, and the development proceeds on a formal basis.

of $1/\alpha$. This can be accomplished either by a formal binomial expansion of the radical in Equation (22) after the square of α has been factored out, or, alternatively, by repeated integration by parts of the numerator of the integrand. In either case there results (see Appendix 2)

$$I(\alpha) \sim (1/\alpha) - I_1 (1/\alpha)^3 + I_2 (1/\alpha)^5 - \dots \quad (26)$$

where

$$I_1 = \frac{1}{2} \left(1 - \frac{1}{2} q \right) \quad (27)$$

and

$$I_2 = (3/4) \left[\int_0^\infty x f dx - q \left(1 - \frac{1}{2} \int_0^\infty f^2 dx \right) + (1/6) q^2 \right] \quad (28)$$

Use of the first term of (26) in (21) results in the familiar

$$\tau \sim 10^{-3} N_0 \cot \theta_0 \quad \text{mrad} \quad (29)$$

which holds at high elevation angles.

Equation (26) suffers from the usual property of an asymptotic expansion — it is not useful at small values of the argument α . There is, it happens, a procedure [10] for converting a divergent series such as Equation (26) into a continued fraction expansion that converges, in this case, for all $\alpha > 0$. The expansion diverges at $\alpha = 0$, however, and converges only slowly when α is near zero. Rather than apply the procedure directly, therefore, the integral $I(\alpha)$ is expanded, with α small, as

$$I(\alpha) = i_0 + i_1 \alpha + \dots \quad (30)$$

where (Appendix 3)

$$i_0 = I(0) = 2 \int_0^\infty \frac{f''}{\sqrt{x + q(1-f)}} \frac{f''}{(1 + q f')^2} dx \quad (31)$$

and

$$i_1 = I'(0) = 2 f'(0) [1 + q f'(0)] \quad (32)$$

Noting that Equation (26) approximates $I(\alpha)$ when α is large, and that Equation (30) approximates $I(\alpha)$ when α is small, the approximation of $I(\alpha)$ over the entire range of α is accomplished by means of a ratio of polynomials in α . The coefficients of the polynomials are chosen in such a way that the expansion of the ratio in inverse powers of α agrees with the leading terms of Equation (26) on the one hand, and its expansion in ascending powers of α agrees with the leading terms of Equation (30) on the other hand. This method of approximation insures accuracy if α is either large or small. Accuracy with intermediate values is obtained by the inclusion of a sufficient number of terms from each series expansion. The number of terms used here — three from Equation (26) and two from Equation (30) — is not necessarily optimum, but worked out well when the method was applied to an exponential profile. It is evident that the method requires a certain degree of smoothness in the profile if an accurate approximation is to be obtained with this number of terms.

4.4 Form of the Approximation

Consider a rational function of α , $F(\alpha; F_1, F_2, f_0, f_1)$, which depends on four parameters F_1, F_2, f_0 and f_1 , and which is expressed in the form of a continued fraction

$$F(\alpha; F_1, F_2; f_0, f_1) = \frac{1}{\alpha + \frac{\underline{f}_1}{\alpha + \frac{\underline{f}_2}{\alpha + \frac{\underline{f}_3}{\alpha + \underline{f}_4}}}} \quad (33)$$

where the intermediate constants $\underline{f}_1, \underline{f}_2, \underline{f}_3$, and \underline{f}_4 are calculated from the set of parameters F_1, F_2, f_0 , and f_1 using in sequence

$$\underline{f}_1 = F_1 \quad (34)$$

$$\underline{f}_2 = (F_2 / \underline{f}_1) - \underline{f}_1 \quad (35)$$

$$\underline{f}_3 = \underline{f}_2 / \left[f_0^2 \underline{f}_1 \left(1 + \frac{\underline{f}_1}{\underline{f}_2} \right) - (1 + f_1 \underline{f}_1) \right] \quad (36)$$

$$\underline{f}_4 = f_0 \underline{f}_1 \underline{f}_3 / \underline{f}_2 \quad (37)$$

On clearing the denominator of Equation (33) of fractions, and expanding the resulting fractional form by long division in descending powers of α

$$F(\alpha; F_1, F_2; f_0, f_1) = (1/\alpha) - F_1 (1/\alpha)^3 + F_2 (1/\alpha)^5 - \text{constant} \cdot (1/\alpha)^7 \dots (38)$$

If the long division is carried out using ascending powers of α

$$F(\alpha; F_1, F_2; f_0, f_1) = f_0 - f_1 \alpha + \text{constant} \alpha^2 \dots (39)$$

Thus the function $F(\alpha; F_1, F_2, f_0, f_1)$ is well-suited to approximate $I(\alpha)$, Equation (21), provided that the parameters F_1, F_2, f_0 , and f_1 are chosen as I_1, I_2, i_0 , and i_1 respectively, i.e., $I(\alpha)$ is approximately equal to $F(\alpha; I_1, I_2; i_0, i_1)$. Note that these latter parameters, Equations (27), (28), (31), and (32), depend only on the refractivity through the parameter q and the profile $f(x)$. Their calculation is independent of satellite position, and numerical integration prior to each satellite pass need not be ruled out. If f is a given model profile, moreover, the integrals can be evaluated analytically if the functional form of $f(x)$ permits, or, otherwise, may be evaluated numerically and curve or surface fitted empirically.

4.5 Formula for Range Error

The range error Equation (13) may be written as the sum of the difference between the electrical and geometric distances along the ray path and of the difference between the geometric distance along the ray path and the slant range

$$\Delta R = (R_e - R_g) + (R_g - R) (40)$$

The first term in Equation (40), the difference along the ray path, is, after division by r_0

$$(R_e - R_g)/r_0 = (1/r_0) \int_0^\infty [10^{-6} N(h)/\sin \theta] dr = \frac{1}{2} 10^{-6} N_0 p J(\alpha) \quad (41)$$

where

$$J(\alpha) = \int_0^\infty \frac{f}{\sqrt{x + \alpha^2 - q(1-f)}} dx \quad (42)$$

While the geometrical difference ($R_g - R$) between the ray and straight-line paths is smaller than the electrical difference ($R_e - R_g$) along the ray path, it is not entirely negligible. In Appendix 4 an expression for this difference is derived. Substituting Equations (4-8) and (41) into (40), the expression for the range error is

$$\Delta R/r_0 = \frac{1}{2} 10^{-6} N_0 p [M(\alpha) - \frac{1}{2} p Q L^2(\alpha)] \quad (43)$$

where

$$M(\alpha) = J(\alpha) + q \left[I(\alpha) - \frac{1}{2} K(\alpha) - \frac{1}{2} \alpha I^2(\alpha) + \frac{1}{12} q I^3(\alpha) \right] \quad (44)$$

with

$$K(\alpha) = \int_0^\infty \frac{-2 f f'}{\sqrt{x + \alpha^2 - q(1-f)}} dx \quad (45)$$

The expansions of $J(\alpha)$ and $K(\alpha)$ for large and for small values of α have been given in Appendices 2 and 3. Thus it is possible to calculate these integrals in the same way as was done for $I(\alpha)$ by using Equation (33) with suitable choices for the parameters F_1 , F_2 , f_0 , and f_1 and then to compute $M(\alpha)$ by substituting

the calculated values of $I(\alpha)$, $J(\alpha)$ and $K(\alpha)$ into Equation (44). It is more efficient to calculate $M(\alpha)$ directly however. Thus, substituting Equations (26), (2-3), and (2-6) into (44)

$$M(\alpha) \sim (1/\alpha) - M_1 (1/\alpha)^3 + M_2 (1/\alpha)^5 - \dots \quad (46)$$

where

$$M_1 = \frac{1}{2} \left[\int_0^\infty x f dx - q \left(1 - \frac{1}{2} \int_0^\infty f^2 dx \right) \right] \quad (47)$$

$$M_2 = \left(\frac{3}{4} \right) \left[\frac{1}{2} \int_0^\infty x^2 f dx - q \left(\frac{1}{6} + \int_0^\infty x f dx - \frac{1}{2} \int_0^\infty x f^2 dx \right) + q^2 \left(\frac{1}{2} - \frac{1}{2} \int_0^\infty f^2 dx + \frac{1}{6} \int_0^\infty f^3 dx \right) \right] \quad (48)$$

Substituting Equations (30), (3-7) and (3-11) into (44)

$$M(\alpha) = m_0 - m_1 \alpha + \dots \quad (49)$$

$$m_0 = j_0 + q i_0 + (1/12) q^2 i_0^3 - \frac{1}{2} q k_0 \quad (50)$$

$$m_1 = j_1 + \frac{1}{2} q i_0^2 \left(1 + \frac{1}{2} q i_1 \right) \quad (51)$$

$M(\alpha)$ is now calculated using Equations (33-37); i.e., is given by $F(\alpha; M_1, M_2, m_0, m_1)$.

A formula for the range-rate correction can be derived (Appendix 5) and, presumably, modelled as above, but this has not been done. In practice (with

the Goddard Range and Range Rate System) the range-rate correction can be obtained by dividing the difference between successive range corrections by the corresponding time interval.

5. CORRECTIONS USING KNOWN ELEVATION ANGLE

The formulas for the corrections ΔE and ΔR as functions of β and R rather than as functions of α are more difficult to obtain Reichly [4] in an excellent treatment of the problem, has used a perturbation method to obtain the first two terms of the expansion of these corrections in powers of the surface refractivity. This, in the notation and with the approximations used here, is equivalent to an expansion in powers of the parameter q which may be as large as 0.64 at $N_0 = 450$. Such an expansion is accurate at large values of the elevation angles as may be determined by an examination of the asymptotic expansions, in which the higher powers of q appear only in the higher order terms. More terms are needed, however, if formulas numerically accurate at small values of the elevation angle are to be obtained by this method. Such formulas can be obtained by expansion, not in powers of q , but in powers of ρ , β and $1/\beta$, and by the use of the method of the preceding paragraphs.

5.1 Elevation Correction

It is assumed that R is large enough to permit ΔE to be calculated from the first two terms of its expansion in powers of the normalized inverse range ρ .

Setting

$$\Delta E = 10^{-3} N_0 (1/p) [U(\beta) - \rho V(\beta)] \cos E \quad \text{mrad} \quad (52)$$

where U and V are functions of β that must be determined, noting that⁴

$$\cos \theta \stackrel{0}{=} \cos E \quad (53)$$

and that

$$\alpha \stackrel{0}{=} \beta + (1/p) \Delta E \cos E = \beta + \frac{1}{2} \rho (U - \rho V) \quad (54)$$

there results, equating Equations (24) and (52)

$$U - \rho V = I \left(\beta + \frac{1}{2} \rho U - \frac{1}{2} \rho^2 V \right) - \rho L \left(\beta + \frac{1}{2} \rho U - \frac{1}{2} \rho^2 V \right) \quad (55)$$

If I and L in Equation (55) above are expanded in powers of ρ , and if coefficients of like powers of ρ are equated

$$U = I \left(\beta + \frac{1}{2} \rho U \right) \quad (56)$$

$$V(\beta) = L \left(\beta + \frac{1}{2} \rho U \right) / \left[1 - \frac{1}{2} \rho I' \left(\beta + \frac{1}{2} \rho U \right) \right] \quad (57)$$

Equation (56) gives U implicitly as a function of β . Taking

$$U(\beta) \sim (1/\beta) - U_1 (1/\beta)^3 + U_2 (1/\beta)^5 - \dots \quad (58)$$

4. $\cos E = \cos(\theta_0 - \Delta E) = \cos \theta_0 (\cos \Delta E - \sin \Delta E \sin \theta_0 / \cos \theta_0)$
 $\stackrel{0}{=} \cos \theta_0 [1 - 10^{-6} N_0 (I - \rho L) \alpha] \stackrel{0}{=} \cos \theta_0$ since $I > \rho L$ and $\alpha I < 1$.

the parameters U_1 and U_2 are determined by substituting Equation (58) into Equation (56) with I in (56) given by Equation (26), expanding in powers of $1/\beta$, and equating coefficients of like powers. The results are

$$U_1 = \frac{1}{2} \left(1 + \frac{1}{2} q \right) \quad (59)$$

$$U_2 = \left(\frac{3}{4} \right) \left[\int_0^x x f dx + q \left(\frac{1}{3} + \frac{1}{2} \int_0^x f^2 dx \right) + (1/6) q^2 \right] \quad (60)$$

At small values of

$$U(\beta) = u_0 - u_1 \beta + u_2 \beta^2 - \dots \quad (61)$$

To obtain u_0 , β is set equal to zero in Equation (56) to obtain

$$u_0 = I \left(\frac{1}{2} q u_0 \right) \quad (62)$$

which can be solved iteratively. The parameters u_1 and u_2 are obtained by differentiating Equation (56)

$$U'(\beta) = I' \left(\beta + \frac{1}{2} q u_0 \right) / \left[1 - \frac{1}{2} q I' \left(\beta + \frac{1}{2} q u_0 \right) \right] \quad (63)$$

whence

$$u_1 = - I' \left(\frac{1}{2} q u_0 \right) / \left[1 - \frac{1}{2} q I' \left(\frac{1}{2} q u_0 \right) \right] \quad (64)$$

Similarly, a second differentiation results in

$$u_2 = I'' \left(\frac{1}{2} q u_0 \right) / \left[1 - \frac{1}{2} q I' \left(\frac{1}{2} q u_0 \right) \right]^3 \quad (65)$$

Using U_1 , U_2 , u_0 , and u_1 , $U(\beta)$ in Equation (52) is approximated by

$$F(\beta; U_1, U_2; u_0, u_1).$$

To calculate $V(\beta)$ Equation (51), Equation (63) is solved for $I'(\beta + 1/2 q u_0)$ and substituted into Equation (57):

$$V = \left[1 - \beta U(\beta) - \frac{1}{4} q U^2(\beta) \right] \left[1 + \frac{1}{2} q U'(\beta) \right] \quad (66)$$

Here $U'(\beta)$ may be calculated using U_1 , U_2 , u_0 , and u_1 as shown in Appendix 6.

5.2 Range Correction

Substitution of Equation (54) into (43) and expansion in powers of ρ gives (using Equation 5-1)

$$\Delta R/r_0 = \frac{1}{2} 10^{-6} N_0 \rho \left[M \left(\beta + \frac{1}{2} q U \right) + \frac{1}{2} Q \rho L^2 \left(\beta + \frac{1}{2} q U \right) \right] \quad (67)$$

which becomes

$$\Delta R/r_0 = \frac{1}{2} 10^{-6} N_0 \rho \left\{ W(\beta) + \frac{1}{2} Q \rho \left[1 - \beta U(\beta) - \frac{1}{4} q U^2(\beta) \right]^2 \right\} \quad (68)$$

where, using the asymptotic expansion of U in $M(\beta + 1/2 q U)$ as given by Equation (46)

$$W(\beta) \sim (1/\beta) - W_1 (1/\beta)^3 + W_2 (1/\beta)^5 - \dots \quad (69)$$

$$W_1 = \frac{1}{2} \left(\int_0^\infty x f dx + \frac{1}{2} q \int_0^\infty f^2 dx \right) \quad (70)$$

$$W_2 = (3/4) \left[\frac{1}{2} \int_0^\infty x^2 f \, dx + (q/6) \left(1 + 3 \int_0^\infty x f^2 \, dx \right) + \frac{1}{6} q^2 \int_0^\infty f^3 \, dx \right] \quad (71)$$

and where, setting $\beta = 0$

$$W(\beta) = w_0 + w_1 \beta + \dots \quad (72)$$

$$w_0 = J \left(\frac{1}{2} q u_0 \right) + q \left[u_0 - \frac{1}{2} K \left(\frac{1}{2} q u_0 \right) - \frac{1}{6} q u_0^3 \right] \quad (73)$$

$$w_1 = 2 \left(1 - \frac{1}{4} q u_0^2 \right) \quad (74)$$

6. EXAMPLE USING AN EXPONENTIAL PROFILE

If an exponential refractivity profile is assumed, the normalized profile f becomes

$$f(x) = e^{-x} \quad (75)$$

The effective height H is equal to the reciprocal of the decay constant of the profile and will be estimated from the refractivity at the station using the empirical formula [10]

$$1/H = \ln \frac{N_0}{N_0 - 7.32 e^{0.005577 N_0}} \quad (76)$$

6.1 Known Arrival-Angle

Most of the required integration, that in Equations (28), (47) and (48), may be performed directly. Equation (3-5) or (31) and Equation (3-12) however, must

be handled numerically. The simplest procedure is to evaluate these integrals numerically [12] at, perhaps, 20 different values of q spread over the expected range $0.1 < q < 0.7$. The values obtained may then be approximated by a polynomial using, since it is simple to apply, a least squares fit. One finds, for example, that a cubic polynomial in q approximates i_0 to better than 1/4 percent over the range 0 to 0.7. As it happened, however, an exponential expression gave a better approximation with fewer empirical constants:

$$i_0 = \sqrt{\pi} (1 - 0.9206 q)^{-0.4468} \pm 0.04\% \quad 0 \leq q \leq 0.7 \quad (77)$$

The integral k_0 was also closely approximated by a similar exponential expression.

The equations for the corrections are collected in Appendix 7.

6.2 Known Elevation Angle

Here, also, the coefficients of the asymptotic expansions, Equations (60), (70), and (71), may be integrated. Equation (62) was solved numerically for u_0 at a number of values of q by Wegstein's method [13], using repeated numerical integration. The calculated values of u_0 were fitted by

$$u_0 = \sqrt{\pi} (1 + 1.4844 q)^{-0.39144} \pm 0.02\% \quad 0 \leq q \leq 0.7 \quad (78)$$

The coefficients u_1 and u_2 were also obtained numerically using the same calculated values of u_0 in Equations (64) and (65) with I' and I'' given by Equations (3-3) and (3-4).

The equations are collected in Appendix 8.

6.3 Numerical Examples and Comparisons

Let the refractivity at the tracking station be $N_0 = 313$. If the tracking station is at sea level, $r_0 = 6369.95$ kilometers. Performing the pre-pass calculations of Appendix 7, $H = 6.951$ km, $p = 0.04672$, $q = 0.2868$, and the equation for the elevation error is

$$\Delta E = 0.313 \cos \theta_0 [i - (6369.95/R) L] \quad \text{mrad} \quad (79)$$

with

$$i = \frac{1}{\sin \theta_0 + \frac{0.0009348}{\sin \theta_0 + \frac{0.002117}{\sin \theta_0 + \frac{0.006054}{\sin \theta_0 + 0.1163}}}} \quad (80)$$

and

$$L = 1 - i \sin \theta_0 + 0.0001565 i^2 \quad (81)$$

The equation for the range correction becomes

$$\Delta R = 0.002176 [m - (913.5/R) L^2 \cos^2 \theta_0] \quad \text{km} \quad (82)$$

$$\begin{aligned}
 m = & \frac{1}{\sin \theta_0 + \frac{0.0008565}{\sin \theta_0 + \frac{0.002173}{\sin \theta_0 + \frac{0.006082}{\sin \theta_0 + \sin \theta_0 + 0.1157}}}} \quad (83)
 \end{aligned}$$

Using twelve values of θ_0 from 0 to 900 milliradians and two different values of the range R (in kilometers) at each value of θ_0 (these ranges are for satellites at heights with respect to the tracking station of 70 km and 475 km), the corrections in Table 1 were calculated. The corrections calculated using a double-precision ray trace program [14] are also listed for comparison.⁶

The largest difference is 0.3 percent. If desired, a final empirical adjustment of the coefficients can be made to reduce this error (Appendix 9). In Table 2 the corrections calculated from Appendix 8, using the values of elevation angle and range listed, are compared with those obtained from the ray-trace program. Since in this case the corrections do not hold for negative elevation angles, all but the smallest of these have been omitted from the table. The maximum error in Table 2, 0.9 percent, is larger than that found in Table 1, apparently because only two terms were retained in Equation (52).

6. These values will be found to differ somewhat from those in the CRPL ray-trace tables [15] which used single precision and slightly different values of r_0 .

Table 1
Calculated Corrections, Arrival Angle Known

PREPASCALCULATIONS1

TIME 0 0 2 (MINUTES:SECONDS:60THS)

CORRECTIONS1

TIME 0 0 2 (MINUTES:SECONDS:60THS)

TABLE1

NO	THETA0 (MRAD)	R (KM)	AE (MRAD)	AE (MRAD) RAY TRACE	AR (KM)	AR (KM) RAY TRACE	PERCENT-ERROR AR	PERCENT-ERROR AR
313	0	1020.2	1.109E01	1.108E01	1.018E01	1.018E01	0.00	-0.06
313	0	2587.1	1.262E01	1.262E01	1.038E01	1.038E01	0.00	-0.06
313	1	1011.3	1.079E01	1.079E01	9.853E02	9.859E02	0.00	-0.06
313	1	2578.2	1.227E01	1.227E01	1.004E01	1.004E01	0.00	-0.06
313	2	1002.3	1.050E01	1.050E01	9.545E02	9.552E02	-0.02	-0.07
313	2	2569.5	1.193E01	1.194E01	9.714E02	9.721E02	-0.01	-0.07
313	4	985.7	9.968E00	9.972E00	8.977E02	8.988E02	-0.04	-0.10
313	4	2552.4	1.131E01	1.131E01	9.122E02	9.131E02	-0.04	-0.10
313	8	953.6	9.031E00	9.041E00	8.003E02	8.016E02	-0.11	-0.16
313	8	2519.0	1.022E01	1.023E01	8.111E02	8.124E02	-0.10	-0.16
313	15	901.8	7.719E00	7.736E00	6.689E02	6.705E02	-0.22	-0.24
313	15	2455.6	8.691E00	8.708E00	6.756E02	6.773E02	-0.19	-0.24
313	30	805.4	5.617E00	5.833E00	4.879E02	4.892E02	-0.28	-0.27
313	30	2360.3	6.498E00	6.513E00	4.908E02	4.921E02	-0.22	-0.27
313	65	633.5	3.589E00	3.594E00	2.900E02	2.904E02	-0.13	-0.16
313	65	2146.8	3.965E00	3.968E00	2.906E02	2.911E02	-0.09	-0.16
313	100	511.9	2.547E00	2.548E00	2.027E02	2.029E02	-0.04	-0.10
313	100	1902.4	2.798E00	2.799E00	2.029E02	2.032E02	-0.02	-0.10
313	200	316.8	1.350E00	1.350E00	1.073E02	1.073E02	0.02	-0.06
313	200	1546.4	1.477E00	1.477E00	1.073E02	1.074E02	0.02	-0.09
313	400	174.9	6.616E01	6.615E01	5.556E03	5.560E03	0.03	-0.08
313	400	1046.4	7.234E01	7.233E01	5.556E03	5.561E03	0.02	-0.09
313	900	89.1	2.234E01	2.233E01	2.774E03	2.776E03	0.03	-0.08
313	900	593.8	2.443E01	2.443E01	2.774E03	2.776E03	0.02	-0.09

NO USED WAS 6369.95

Table 2
Calculated Corrections, Elevation Angle Known

PREPASS CALCULATIONS2
TIME 0 0 4 (MINUTES:SECONDS:60THS)
CORRECTIONS2
TIME 0 0 1 (MINUTES:SECONDS:60THS)

TABLE2

NO	E (MRAD)	R (KM)	ΔE (MRAD)	ΔE (MRAD) RAY TRACE	ΔR (KM) RAY TRACE	ΔR (KM) RAY TRACE	PERCENT-ERROR ΔE	PERCENT-ERROR ΔR
313	1.04	953.6	9.124E00	9.041E00	8.603E-02	8.016E-02	0.92	-0.17
313	2.23	2519.6	1.024E01	1.023E01	8.117E-02	8.124E-02	0.13	-0.09
313	7.26	901.8	7.780E00	7.736E00	6.691E-02	6.705E-02	0.56	-0.21
313	6.29	2465.6	8.710E00	8.708E00	6.763E-02	6.773E-02	0.02	-0.14
313	24.17	805.4	5.835E00	5.833E00	4.876E-02	4.892E-02	0.04	-0.33
313	23.49	2360.3	6.498E00	6.513E00	4.905E-02	4.921E-02	-0.22	-0.31
313	61.41	633.5	3.592E00	3.594E00	2.898E-02	2.904E-02	-0.05	-0.21
313	61.03	2146.8	3.964E00	3.966E00	2.904E-02	2.911E-02	-0.11	-0.21
313	97.45	511.9	2.548E00	2.548E00	2.027E-02	2.029E-02	0.00	-0.14
313	97.20	1962.4	2.798E00	2.799E00	2.029E-02	2.032E-02	-0.02	-0.14
313	198.65	316.8	1.351E00	1.350E00	1.072E-02	1.073E-02	0.03	-0.11
313	198.52	1546.4	1.477E00	1.477E00	1.073E-02	1.074E-02	0.02	-0.12
313	399.34	174.9	6.617E-01	6.615E-01	5.554E-03	5.560E-03	0.03	-0.11
313	399.28	1046.4	7.234E-01	7.233E-01	5.554E-03	5.561E-03	0.02	-0.12
313	899.78	89.1	2.234E-01	2.233E-01	2.773E-03	2.776E-03	0.03	-0.11
313	899.76	593.8	2.443E-01	2.443E-01	2.773E-03	2.776E-03	0.02	-0.12

RO USED WAS 6369.95

Similar checks were made at $N_0 = 200$ and $N_0 = 450$ for both kinds of corrections with equivalent results.

The computer time required to perform the calculations is also shown in the tables. The pre-pass calculations in Table 1 took two sixtieths of a second (IBM 360/95). The two sixtieths of a second used to compute the twenty four pairs of corrections in Table 1 represents a rate of about 700 per second. Programs designed for operational use should show improvements over these times.

7. SUMMARY

A method is given for deriving atmospheric elevation-error and range-error correction equations in a form suitable for use in the processing of satellite tracking data. The method can be applied to any sufficiently-smooth spherically-symmetric model of the atmospheric refractivity. The method was tested by application to an exponential profile and comparison of the resulting corrections with those given by a double-precision ray trace program. The results were in agreement to better than one percent over the entire range of elevation angle ($0 - 90^\circ$) and better than 0.3 percent over most of the range.

An inspection of the coefficients in the correction equations shows that the corrections at high elevations do not depend on the final detail of the refractivity

profile $M(h)$, but rather on the value of the surface refractivity and, in apparent order of importance, the height integrals $\int Ndh$, $\int hNdh$, and $\int N^2dh$, etc. The derivative $N'(0)$ of the refractivity at the tracking station is also significant because it weighs heavily in the determination of both coefficients of the small-angle expansion of the bending integral. Where improved accuracy is required, emphasis should be placed on studies leading to better estimates for these significant quantities.

8. ACKNOWLEDGEMENT

The author wishes to acknowledge the help and co-operation of Dr. S. Ranganwamy, who supplied the ray-trace data used in Tables 1 and 2.

REFERENCES

1. Weisbrod, S. and Anderson, L., "Simple Methods for Computing Tropospheric and Ionospheric Refraction Effects on Radio Waves," Proc. IRE vol. 47, pp. 1770-1777; Oct. 1959.
2. Thayer, G., "A Formula for Radio Ray Refraction in an Exponential Atmosphere," Jour. of Res. of the N.B.S. - D. Radio Propagation Vol. 65, No. 2, March-April 1961, pp. 181-182.
3. Freeman, J., "Range-Error Compensation for a Troposphere With Exponentially Varying Refractivity," Jour. of Res. of the N.B.S. - D. Radio Propagation Vol. 66D, No. 6, Nov.-Dec. 1962, pp. 695-697.

4. Reichly, P., Jet Propulsion Laboratory Space Programs Summary 37-43
vol. IV, Feb. 28, 1967, pp. 314-321.
5. Hopfield, H., "Two-Quartic Tropospheric Refractivity Profile for Correcting Satellite Data," Jour. of Geophysical Research, Vol. 74, No. 18, August 20, 1969, pp. 4487-4499.
6. Rowlandson, L. and Moldt, R., "Derivation of Closed Functions to Compensate Range and Angle Errors in an Exponential Atmosphere," Radio Science, Vol. 4, No. 10, pp. 927-933, October 1969.
7. "Design Evaluation Report, Goddard Range and Range Rate System," General Dynamics Electronics Division Report R-67-042, December 13, 1967. Prepared for Goddard Space Flight Center under contract NAS 5-10555.
8. Berbert, J. and Parker, H., "GEOS Satellite Tracking Corrections for Refraction in the Troposphere," Goddard Space Flight Center Report X-514-70-55, Feb. 1970.
9. Bean, B. and Thayer, G., "Models of the Atmospheric Radio Refractive Index," Proc. IRE vol. 47, May 1959, pp. 740-755. See also Bean, B., and Dutton E., "Radio Meteorology," NBS Monograph 92, March 1, 1966.
10. Wall, H. S., "Analytic Theory of Continued Fractions," D. Van Nostrand Company, Inc., New York, 1948; p. 362.
11. Saxton, J. (ed.), "Advances in Radio Research," Vol. I, Academic Press, N.Y. 1964, p. 65.
12. Krylov, V. I., "Approximate Calculation of Integrals," McMillan, N.Y., 1962.

13. Lance, G. C., "Numerical Methods for High Speed Computers," Iliffe, London, 1960, pp. 134-138.
14. Developed by S. Rangaswamy, National Academy of Science Research Associate, Goddard Space Flight Center.
15. Bean, B. and Thayer, G., "CRPL Exponential Reference Atmosphere," NBS Monograph 4, October 29, 1959.
16. APL/360 Users Manual, International Business Machines Corporation, 1968.

APPENDIX 1

EXPRESSION FOR δ IN TERMS OF τ

The equation for the angle δ follows from the geometry of Figure 2, and from Snell's law for a spherically stratified medium. In Figure 2 r_1 , r_0 , and R are all projected on a line perpendicular to the tangent to the ray path at the satellite. It follows that

$$r_1 \cos \theta_1 = r_0 \cos (\theta_0 - \tau) + R \sin \delta \quad (1-1)$$

Snell's law for a spherically stratified medium is

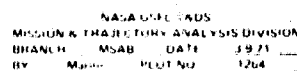
$$n r \cos \theta = n_0 r_0 \cos \theta_0 \quad (1-2)$$

At the high altitude of the satellite n is equal to unity, $r = r_1$, and $\theta = \theta_1$. Combining Equations (1-1) and (1-2) gives

$$\sin \delta = (r_0/R) [n_0 \cos \theta_0 - \cos (\theta_0 - \tau)] \quad (1-3)$$

Approximating $\sin \delta$ by δ , $\sin \tau$ by τ , and $\cos \tau$ by $1 - 1/2 \tau^2$,

$$\delta \approx (r_0/R) \left[\left(10^{-6} N_0 + \frac{1}{2} \tau^2 \right) \cos \theta_0 - \tau \sin \theta_0 \right] \quad (1-4)$$



33

APPENDIX 2

ASYMPTOTIC EXPANSION OF I, J, AND K

Setting $Q = q$ in Equation (22), factoring α out of the radical, and expanding the latter by the binomial theorem

$$I(\alpha) \sim (1/\alpha) \int_0^\infty -f' \left\{ 1 - \frac{1}{2} [x - q(1-f)] (1/\alpha)^2 \right. \\ \left. + (3/8) [x - q(1-f)]^2 (1/\alpha)^4 - \dots \right\} dx \quad (2-1)$$

The terms in (2-1) that do not contain x as a factor may be integrated directly. Those that do contain x can be integrated by parts on the factor containing f' and powers of f . The result of these integrations is given in Equations (26), (27), and (28). (The same result can be obtained by repeatedly integrating the numerators starting with Equation (22), using the definition

$$f^{(-1)} = - \int_x^\infty f \, dx$$

and assuming that $f^{(-1)}$ vanishes with sufficient rapidity at large values of x).

The integral $J(\alpha)$ Equation (42) may be expanded in the same way with the result

$$J(\alpha) \sim (1/\alpha) - J_1 (1/\alpha)^3 + J_2 (1/\alpha)^5 - \dots \quad (2-3)$$

$$J_1 = \frac{1}{2} \left[\int_0^\infty x f \, dx - q \left(1 - \int_0^\infty f^2 \, dx \right) \right] \quad (2-4)$$

$$J_2 = (3/4) \left[\frac{1}{2} \int_0^\infty x^2 f \, dx - q \left(\int_0^\infty x f \, dx - \int_0^\infty x f^2 \, dx \right) \right. \\ \left. + q^2 \left(\frac{1}{2} - \int_0^\infty f^2 \, dx + \frac{1}{2} \int_0^\infty f^3 \, dx \right) \right] \quad (2-5)$$

Similarly Equation (45) yields

$$K(a) \sim (1/a) - K_1 (1/a)^3 + K_2 (1/a)^5 - \dots \quad (2-6)$$

$$K_1 = \frac{1}{2} \int_0^\infty f^2 \, dx - (1/6) q \quad (2-7)$$

$$K_2 = (3/4) \left\{ \int_0^\infty x f^2 \, dx - q \left[\int_0^\infty f^2 \, dx - (2/3) \int_0^\infty f^3 \, dx \right] + (1/12) q^2 \right\} \quad (2-8)$$

APPENDIX 3

INTEGRALS FOR I, J AND K, AND

EXPANSIONS FOR SMALL VALUES OF α

The integral Equation (22) can be integrated by parts. Multiplying the numerator and denominator of the integrand by $1 + qf'$ and integrating the radical by parts

$$I(\alpha) = \frac{2 f'(0)}{1 + q f'(0)} \alpha + 2 \int_0^{\infty} \frac{\sqrt{x + \alpha^2 - q(1-f)} f''}{(1 + q f')^2} dx \quad (3-1)$$

The first term may be absorbed into the integral

$$I(\alpha) = 2 \int_0^{\infty} \frac{[\sqrt{x + \alpha^2 - q(1-f)} - \alpha] f''}{(1 + q f')^2} dx \quad (3-2)$$

Differentiating Equation (3-2)

$$I'(\alpha) = -2 \int_0^{\infty} \left[1 - \frac{\alpha}{\sqrt{x + \alpha^2 - q(1-f)}} \right] \frac{f''}{(1 + q f')^2} dx \quad (3-3)$$

Differentiating Equation (3-3)

$$I''(\alpha) = 2 \int_0^{\infty} \frac{x - q(1-f)}{[x + \alpha^2 - q(1-f)]^{3/2}} \frac{f''}{(1 + q f')^2} dx \quad (3-4)$$

Equation (31) for $I(0)$ is obtained by setting α equal to zero in Equation (3-2).

The integral

$$I(0) = \int_0^{\infty} \frac{-f'}{\sqrt{x - q(1-f)}} dx \quad (3-5)$$

obtained directly from Equation (22) can also be used, but more care must then be exercised since the integrand diverges at $x = 0$.

Equation (32) follows from Equation (3-3) by setting $\alpha = 0$ and integrating. In the same way, from Equation (42)

$$J(\alpha) = -2 \int_0^{\infty} [\sqrt{x + \alpha^2 - q(1-f)} - \alpha] \frac{f' + q(f'^2 - f f'')}{(1 + q f')^2} dx \quad (3-6)$$

$$J(\alpha) = j_0 - j_1 \alpha + \dots \quad (3-7)$$

with

$$j_0 = J(0) = -2 \int_0^{\infty} \sqrt{x - q(1-f)} \frac{f' + q(f'^2 - f f'')}{(1 + q f')^2} dx \quad (3-8)$$

$$j_1 = -2/[1 + q f'(0)] = i_1 / -f'(0) \quad (3-9)$$

Similarly,

$$K(\alpha) = 4 \int_0^{\infty} [\sqrt{x + \alpha^2 - q(1-f)} - \alpha] \frac{f f'' + f'^2 + q f'^3}{(1 + q f')^2} dx \quad (3-10)$$

and setting

$$K(\alpha) = k_0 - k_1 \alpha + \dots \quad (3-11)$$

$$k_0 = 4 \int_0^\infty \sqrt{x - q(1-f)} \frac{f f'' + f'^2 + q f'^3}{(1 + q f')^2} dx \quad (3-12)$$

$$k_1 = -4 f'(0) / [1 + q f'(0)] = 2 i_1 \quad (3-13)$$

APPENDIX 4

APPROXIMATION FOR THE GEOMETRICAL DIFFERENCE

Using Equation (1-2), R_g equation (12) may be written as

$$R_g = \int_{r_0}^{r_1} \frac{(r + n_0^2 r_0^2 \cos^2 \theta_0 n' / n^3) - n_0^2 r_0^2 \cos^2 \theta_0 n' / n^3}{\sqrt{r^2 - (n_0^2 r_0^2 \cos^2 \theta_0) / n^2}} dr \quad (4-1)$$

The left-hand term in the integrand can be integrated. On setting $n(r_1) = 1$ and making use of the geometrical relationship

$$r_1 \sin \theta_1 = R \cos \delta + r_0 \sin (\theta_0 - \tau) \quad (4-2)$$

which follows from Figure 2, and using the exact integral for τ [9]

$$\tau = -r_0 \cos \theta_0 \int_0^\infty \frac{n_0 n'}{n^2 \sqrt{r^2 - n_0^2 r_0^2 \cos^2 \theta_0 / n^2}} dr \quad (4-3)$$

there results

$$(R_g - R) / r_0 = - (R / r_0) (1 - \cos \delta) - (1 - \cos \tau) \sin \theta_0 \quad (4-4)$$

$$+ (\tau - \sin \tau) \cos \theta_0 + \cos^2 \theta_0 \int_0^\infty \frac{n_0 (n - n_0) n'}{n^3 \sqrt{r^2 - \frac{n_0^2 r_0^2 \cos^2 \theta_0}{n^2}}} dr$$

On setting

$$\cos \vartheta \approx 1 - \frac{1}{2} \vartheta^2 \quad (4-5)$$

$$\sin \tau \approx \tau - \tau^3/6 \quad (4-6)$$

$$\cos \tau \approx 1 - \frac{1}{2} \tau^2 \quad (4-7)$$

and making the usual approximations in the integrals, there results

$$\begin{aligned} (P_g - R)/r_0 = \frac{1}{2} 10^{-6} N_0 p_0 \left[\left(I - \frac{1}{2} K - \frac{1}{2} \alpha I^2 + q I^3/12 \right) \right. \\ \left. - (r_0 p/2 R) \left(1 - \alpha I + \frac{1}{4} q I^2 \right)^2 \right] \quad (4-8) \end{aligned}$$

APPENDIX 5

RANGE-RATE CORRECTION

The range-rate correction is the difference between $n(r_1)$ times the projection of the satellite velocity on the ray path at the satellite and its projection on the straight line joining the satellite and the ground station.

It may be found by differentiating Equation (44) with respect to time. Using primes to denote differentiation with respect to α , and the dot notation to denote differentiation with respect to time, and making use of the relation (obtained by performing the indicated differentiation and then integrating by parts)

$$J'(\alpha) - \frac{1}{2} q K'(\alpha) = -2 [1 - \alpha I(\alpha)] \quad (5-1)$$

one obtains

$$\begin{aligned} \Delta \dot{R}/r_0 = (\dot{R}_e - \dot{R})/r_0 = & - \dot{\theta}_0 10^{-6} N_0 \cos \theta_0 L(\alpha) \left\{ 1 - \frac{1}{2} q I'(\alpha) \right. \\ & \left. - \rho \frac{1}{2} q \left[I(\alpha) + \alpha I'(\alpha) - \frac{1}{2} q I(\alpha) I'(\alpha) \right] + \frac{1}{2} \dot{\theta}^2 \dot{R}/r_0 \right\} \end{aligned} \quad (5-2)$$

APPENDIX 6

CALCULATION OF $U'(\beta)$

From Equations (59) and (60)

$$U' \sim - [(1/\beta)^2 - 3 U_1 (1/\beta)^4 + 5 U_2 (1/\beta)^6 - \dots] \quad (6-1)$$

$$U' \sim - [u_1 - 2 u_2 \beta + \dots] \quad (6-2)$$

The calculation of U' is accomplished by means of the form

$$G(\beta) = \frac{1}{\beta^2 + \frac{g_1}{1 + \frac{g_2}{\beta^2 + g_4 \beta + g_3}}} \quad (6-3)$$

where

$$g_1 = G_1 \quad (6-4)$$

$$g_2 = (G_2/G_1) - g_1 \quad (6-5)$$

$$g_3 = g_2/(g_1 g_0 - 1) \quad (6-6)$$

$$g_4 = g_3^2 g_1 g_1 / g_2 \quad (6-7)$$

whence

$$G = (1/\beta)^2 - G_1 (1/\beta)^4 + G_2 (1/\beta)^6 - \text{const.}/\beta^7 \dots \quad (6-8)$$

and

$$G = g_0 - g_1 \beta + \text{constant } \beta^2 + \dots \quad (6-9)$$

Setting $G_1 = 3U_1$, $G_2 = 5U_2$, $g_0 = u_1$, and $g_1 = 2u_2$, U' will be approximately equal to minus G .

APPENDIX 7

CORRECTION EQUATIONS FOR AN EXPONENTIAL PROFILE, ARRIVAL ANGLE KNOWN

$$H = 1/\ln \frac{N_0}{N_0 - 7.32 e^{0.005577 N_0}} \quad \text{km}$$

$$p = \sqrt{2H/r_0}$$

$$q = 10^{-6} N_0 r_0 / H$$

$$i_0 = \sqrt{\pi} (1 - 0.9206 q)^{-0.4468}$$

$$i_1 = 2/(1 - q)$$

$$I_1 = \frac{1}{2} \left(1 - \frac{1}{2} q \right)$$

$$I_2 = 0.75 [1 - 0.75 q + (1/6) q^2]$$

$$k_0 = \sqrt{2\pi} (1 - 0.9408 q)^{-0.4759}$$

$$m_0 = i_0 \left(1 + q + \frac{1}{12} q^2 i_0^2 \right) - \frac{1}{2} q k_0$$

$$m_1 = 2 \left(1 + \frac{1}{4} q i_0^2 \right) / (1 - q)$$

$$M_1 = \frac{1}{2} \left(1 - \frac{3}{4} q \right)$$

$$M_2 = 0.75 \left(1 - \frac{25}{24} q + \frac{11}{36} q^2 \right)$$

$$i = F(\sin \theta_0; p^2 I_1, p^4 I_2; i_0/p, i_1/p^2)$$

$$L = 1 - i \sin \theta_0 + \frac{1}{2} 10^{-6} N_0 i^2$$

$$\Delta E = 0.001 N_0 \cos^2 \theta_0 (i - r_0 L/R) \quad \text{mrad}$$

$$m = F(\sin^2 \theta_0; p^2 M_1, p^4 M_2; m_0/p, m_1/p^2)$$

$$\Delta R = 10^{-6} N_0 H \left(m - \frac{1}{2} 10^{-6} N_0 r_0^2 \cos^2 \theta_0 L^2/RH \right) \quad \text{km}$$

The function i above is equal to $I(\alpha)/p$, and the intermediate constant, Equations (34-37), associated with i can be obtained from those of $I(\alpha)$ by multiplying the first three by p^2 and the last by p .

The programs used to perform the calculations were written in APL language [16], and are shown in Figure 3.

```

VPREPASSCALCULATIONS1[[]]V
V PREPASSCALCULATIONS1
[1] HH+10NNO+NNO-7.32*0.005577*NNO
[2] PSQUARED<2*HH+R0
[3] P+PSQUARED*0.5
[4] Q<1E-6*NNO*R0+HH
[5] F0+I0+((O1)*0.5)*(1-0.9206*Q)*-0.4468
[6] F1+I1+2*1-Q
[7] FF1+II1+0.5*1-0.5*Q
[8] FF2+II2+0.75*1+Q*-0.75+Q+6
[9] CALCULATEF
[10] I1+PSQUARED*F1
[11] I2+PSQUARED*F2
[12] I3+PSQUARED*F3
[13] I4+P*F4
[14] K0+((O2)*0.5)*(1-0.9408*Q)*-0.4759
[15] F0+M0+((I0*1+Q*1+Q*I0*I0+12)-Q*K0+2
[16] F1+M1+2*(1+Q*I0*I0+4)*1-Q
[17] FF1+MM1+0.5*1-0.75*Q
[18] FF2+MM2+0.75*1+Q*(-25+24)+11*Q+36
[19] CALCULATEF
[20] M1+PSQUARED*F1
[21] M2+PSQUARED*F2
[22] M3+PSQUARED*F3
[23] M4+P*F4
V
VCORRECTIONS1[[]]V
V CORRECTIONS1
[1] SIN+10THETA0*0.001
[2] COS+20THETA0*0.001
[3] I<SIN+I1/SIN+I2/SIN+I3/SIN+I4
[4] LL+1-I*SIN-0.5*1E-6*FF0*I
[5] DLE+0.001*NNO*COS*I-R0*LL+RR
[6] M<SIN+M1/SIN+M2/SIN+M3/SIN+M4
[7] DRR+1E-6*NNO*HH*M-0.5*1E-6*NNO*R0*R0*COS*COS*LL*LL+RR*HH
V
VCALCULATEF[[]]V
V CALCULATEF
[1] F1+FF1
[2] F2+(FF2/F1)-F1
[3] F3+F2/(F0*F0*F1*1+F1/F2)-1+F1*F1
[4] F4+F0*F1*F3/F2
V

```

Figure 3. APL Programs for Corrections, Arrival Angle Known.

APPENDIX 8

CORRECTION EQUATIONS FOR AN EXPONENTIAL PROFILE, ELEVATION ANGLE KNOWN

H, p, and q are calculated as in Appendix 7.

$$I' \left(\frac{1}{2} q u_0 \right) = -2 (1 + 1.482 q)^{-0.3826}$$

$$I'' \left(\frac{1}{2} q u_0 \right) = 2 \sqrt{\pi} (1 + 1.71 q)^{0.1}$$

$$u_0 = \sqrt{\pi} (1 + 1.4844 q)^{-0.39144}$$

$$u_1 = -I' \left(\frac{1}{2} q u_0 \right) / \left[1 - \frac{1}{2} q I' \left(\frac{1}{2} q u_0 \right) \right]$$

$$u_2 = I'' \left(\frac{1}{2} q u_0 \right) / \left[1 - \frac{1}{2} q I' \left(\frac{1}{2} q u_0 \right) \right]^3$$

$$U_1 = \frac{1}{2} \left(1 + \frac{1}{2} q \right)$$

$$U_2 = \frac{3}{4} \left(1 + \frac{7}{12} q + \frac{1}{6} q^2 \right)$$

$$K \left(\frac{1}{2} q u_0 \right) = \sqrt{2\pi} (1 + 1.6454 q)^{-0.583}$$

$$w_0 = u_0 \left(1 + q - \frac{1}{6} q u_0^2 \right) - \frac{1}{2} q K \left(\frac{1}{2} q u_0 \right)$$

$$w_1 = 2 \left(1 - \frac{1}{4} q u_0^2 \right)$$

$$W_1 = \frac{1}{2} \left(1 + \frac{1}{4} q \right)$$

$$W_2 = \frac{3}{4} \left(1 + \frac{7}{24} q + \frac{1}{18} q^2 \right)$$

$$u = F(\sin E; p^2 U_1, p^4 U_2; u_0/p, u_1/p^2)$$

$$u' = -G(\sin E; 3 p^2 U_1, 5 p^4 U_2; u_1/p^2, 2 u_2/p^3)$$

$$v = \left(1 - u \sin E - \frac{1}{2} 10^{-6} N_0 u^2\right) (1 + 10^{-6} N_0 u')$$

$$\Delta E = 0.001 N_0 \cos E (u - v r_0/R) \quad \text{mrad}$$

$$w = F(\sin e; p^2 W_1, p^4 W_2; w_0/p, w_1/p^2)$$

$$\Delta R = 10^{-6} N_0 H \left[w - \frac{1}{2} 10^{-6} N_0 \frac{r_0^2}{RH} \cos^2 E \left(1 - u \sin E - \frac{1}{2} 10^{-6} N_0 u^2\right)^2 \right] \quad \text{km}$$

The corresponding APL program is shown in Figure 4.

```

VPRSPASSCALCULATIONS2[U]V
▽ PREPASSCALCULATIONS2
[1] HH←+0.0000000-7.32*0.005577*PPO
[2] PSQUARED←2*HH*RO
[3] P←PSQUARED*0.5
[4] Q←1E-6*NN0*RO*HH
[5] IFU←2*(1+1.482*Q)*-0.3826
[6] DEN←1-0.5*Q*IPU
[7] IPPU←2*((01)*0.5)*(1+1.71*Q)*0.1
[8] F0+U0+((01)*0.5)*(1+1.4844*Q)*-0.39144
[9] F1+U1+-IPU:DEN
[10] FF1+UU1+0.5*1+0.5*Q
[11] FF2+UU2+0.75*1+(Q*(7÷12))+Q÷6
[12] CALCULATEE
[13] U1←PSQUARED*E1
[14] U2←PSQUARED*E2
[15] U3←PSQUARED*E3
[16] U4←P*E4
[17] G0←U0+U1
[18] G1←UP1+U2+IPPU:DEN*3
[19] GG1←UUP1+3*UU1
[20] GG2←UUP2+5*UU2
[21] CALCULATEE
[22] UP1←PSQUARED*Q1
[23] UP2←PSQUARED*Q2
[24] UP3←PSQUARED*Q3
[25] UP4←P*Q4
[26] KU←((02)*0.5)*(1+1.8454*Q)*-0.583
[27] F0+W0+(U0*1+Q*1-Q*U0*U0÷6)-0.5*Q*KU
[28] F1+W1+2*1-0.25*Q*U0*U0
[29] FF1+WW1+0.5*1+0.25*Q
[30] FF2+WW2+0.75*1+Q*(7÷24)+Q÷18
[31] CALCULATEE
[32] W1←PSQUARED*E1
[33] W2←PSQUARED*E2
[34] W3←PSQUARED*E3
[35] W4←P*E4
▽
VCORRECTIONS2[Q]V
▽ CORRECTIONS2
[1] SINE←10EE*0.001
[2] COSE←20EE*0.001
[3] U←SINE+U1÷SINE+U2÷SINE+U3÷SINE+U4
[4] UP←-(SINE*SINE)+UP1÷1+UP2÷1+UP3÷SINE*UP4+SINE
[5] PAREN←1-U*SINE+0.5*1E-6*NN0*U
[6] DEE←0.001*NN0*COSE*U-PAREN*(1+1E-6*NN0*UP)*RO+RP
[7] W←SINE+W1÷SINE+W2÷SINE+W3÷SINE+W4
[8] DRR←1E-6*NN0*HH*W+0.5*1E-6*NN0*(FH*RF)*(RO*COSE*PAREN)*2
▽
VCALCULATEG[B]V
▽ CALCULATEG
[1] G1←GG1
[2] G2←(GG2÷G1)-G1
[3] G3←G2÷1+G1*G0
[4] G4←G3*G3*G1*G1+G2
▽

```

Figure 4. APL Programs for Corrections, Elevation Angle Known.

APPENDIX 9

IMPROVED FIT

The correction equations given in Appendix 7 can easily be modified to give a closer fit to ray-trace calculations. The method chosen was to modify the coefficients \underline{f}_2 , \underline{f}_3 , and \underline{f}_4 using a least squares adjustment that brought $F(\alpha; 1/2, 3/4; \sqrt{\pi}, 2)$, (which is the approximation for $I(\alpha)$ where $q = 0$) into better agreement with its theoretical value $\sqrt{\pi} e^{\alpha^2} \operatorname{erfc} \alpha$. The factors by which \underline{f}_2 , \underline{f}_3 , and \underline{f}_4 are multiplied are 1.08885, 1.320903, and 1.21313 respectively. The resulting corrections are shown in Table 3.

Table 3
Calculated Corrections, F Modified

V CALCULATEE(U) V CALCULATEE		PREPARED CALCULATIONS CORRECTIONS		TABLE 1		TABLE 2		TABLE 3		TABLE 4		TABLE 5		TABLE 6		TABLE 7		TABLE 8		TABLE 9		TABLE 10		TABLE 11		TABLE 12		TABLE 13		TABLE 14		TABLE 15		TABLE 16		TABLE 17		TABLE 18		TABLE 19		TABLE 20		TABLE 21		TABLE 22		TABLE 23		TABLE 24		TABLE 25		TABLE 26		TABLE 27		TABLE 28		TABLE 29		TABLE 30		TABLE 31		TABLE 32		TABLE 33		TABLE 34		TABLE 35		TABLE 36		TABLE 37		TABLE 38		TABLE 39		TABLE 40		TABLE 41		TABLE 42		TABLE 43		TABLE 44		TABLE 45		TABLE 46		TABLE 47		TABLE 48		TABLE 49		TABLE 50		TABLE 51		TABLE 52		TABLE 53		TABLE 54		TABLE 55		TABLE 56		TABLE 57		TABLE 58		TABLE 59		TABLE 60		TABLE 61		TABLE 62		TABLE 63		TABLE 64		TABLE 65		TABLE 66		TABLE 67		TABLE 68		TABLE 69		TABLE 70		TABLE 71		TABLE 72		TABLE 73		TABLE 74		TABLE 75		TABLE 76		TABLE 77		TABLE 78		TABLE 79		TABLE 80		TABLE 81		TABLE 82		TABLE 83		TABLE 84		TABLE 85		TABLE 86		TABLE 87		TABLE 88		TABLE 89		TABLE 90		TABLE 91		TABLE 92		TABLE 93		TABLE 94		TABLE 95		TABLE 96		TABLE 97		TABLE 98		TABLE 99		TABLE 100		TABLE 101		TABLE 102		TABLE 103		TABLE 104		TABLE 105		TABLE 106		TABLE 107		TABLE 108		TABLE 109		TABLE 110		TABLE 111		TABLE 112		TABLE 113		TABLE 114		TABLE 115		TABLE 116		TABLE 117		TABLE 118		TABLE 119		TABLE 120		TABLE 121		TABLE 122		TABLE 123		TABLE 124		TABLE 125		TABLE 126		TABLE 127		TABLE 128		TABLE 129		TABLE 130		TABLE 131		TABLE 132		TABLE 133		TABLE 134		TABLE 135		TABLE 136		TABLE 137		TABLE 138		TABLE 139		TABLE 140		TABLE 141		TABLE 142		TABLE 143		TABLE 144		TABLE 145		TABLE 146		TABLE 147		TABLE 148		TABLE 149		TABLE 150		TABLE 151		TABLE 152		TABLE 153		TABLE 154		TABLE 155		TABLE 156		TABLE 157		TABLE 158		TABLE 159		TABLE 160		TABLE 161		TABLE 162		TABLE 163		TABLE 164		TABLE 165		TABLE 166		TABLE 167		TABLE 168		TABLE 169		TABLE 170		TABLE 171		TABLE 172		TABLE 173		TABLE 174		TABLE 175		TABLE 176		TABLE 177		TABLE 178		TABLE 179		TABLE 180		TABLE 181		TABLE 182		TABLE 183		TABLE 184		TABLE 185		TABLE 186		TABLE 187		TABLE 188		TABLE 189		TABLE 190		TABLE 191		TABLE 192		TABLE 193		TABLE 194		TABLE 195		TABLE 196		TABLE 197		TABLE 198		TABLE 199		TABLE 200		TABLE 201		TABLE 202		TABLE 203		TABLE 204		TABLE 205		TABLE 206		TABLE 207		TABLE 208		TABLE 209		TABLE 210		TABLE 211		TABLE 212		TABLE 213		TABLE 214		TABLE 215		TABLE 216		TABLE 217		TABLE 218		TABLE 219		TABLE 220		TABLE 221		TABLE 222		TABLE 223		TABLE 224		TABLE 225		TABLE 226		TABLE 227		TABLE 228		TABLE 229		TABLE 230		TABLE 231		TABLE 232		TABLE 233		TABLE 234		TABLE 235		TABLE 236		TABLE 237		TABLE 238		TABLE 239		TABLE 240		TABLE 241		TABLE 242		TABLE 243		TABLE 244		TABLE 245		TABLE 246		TABLE 247		TABLE 248		TABLE 249		TABLE 250		TABLE 251		TABLE 252		TABLE 253		TABLE 254		TABLE 255		TABLE 256		TABLE 257		TABLE 258		TABLE 259		TABLE 260		TABLE 261		TABLE 262		TABLE 263		TABLE 264		TABLE 265		TABLE 266		TABLE 267		TABLE 268		TABLE 269		TABLE 270		TABLE 271		TABLE 272		TABLE 273		TABLE 274		TABLE 275		TABLE 276		TABLE 277		TABLE 278		TABLE 279		TABLE 280		TABLE 281		TABLE 282		TABLE 283		TABLE 284		TABLE 285		TABLE 286		TABLE 287		TABLE 288		TABLE 289		TABLE 290		TABLE 291		TABLE 292		TABLE 293		TABLE 294		TABLE 295		TABLE 296		TABLE 297		TABLE 298		TABLE 299		TABLE 300		TABLE 301		TABLE 302		TABLE 303		TABLE 304		TABLE 305		TABLE 306		TABLE 307		TABLE 308		TABLE 309		TABLE 310		TABLE 311		TABLE 312		TABLE 313		TABLE 314		TABLE 315		TABLE 316		TABLE 317		TABLE 318		TABLE 319		TABLE 320		TABLE 321		TABLE 322		TABLE 323		TABLE 324		TABLE 325		TABLE 326		TABLE 327		TABLE 328		TABLE 329		TABLE 330		TABLE 331		TABLE 332		TABLE 333		TABLE 334		TABLE 335		TABLE 336		TABLE 337		TABLE 338		TABLE 339		TABLE 340		TABLE 341		TABLE 342		TABLE 343		TABLE 344		TABLE 345		TABLE 346		TABLE 347		TABLE 348		TABLE 349		TABLE 350		TABLE 351		TABLE 352		TABLE 353		TABLE 354		TABLE 355		TABLE 356		TABLE 357		TABLE 358		TABLE 359		TABLE 360		TABLE 361		TABLE 362		TABLE 363		TABLE 364		TABLE 365		TABLE 366		TABLE 367		TABLE 368		TABLE 369		TABLE 370		TABLE 371		TABLE 372		TABLE 373		TABLE 374		TABLE 375		TABLE 376		TABLE 377		TABLE 378		TABLE 379		TABLE 380		TABLE 381		TABLE 382		TABLE 383		TABLE 384		TABLE 385		TABLE 386		TABLE 387		TABLE 388		TABLE 389		TABLE 390		TABLE 391		TABLE 392		TABLE 393		TABLE 394		TABLE 395		TABLE 396		TABLE 397		TABLE 398		TABLE 399		TABLE 400		TABLE 401		TABLE 402		TABLE 403		TABLE 404		TABLE 405		TABLE 406		TABLE 407		TABLE 408		TABLE 409		TABLE 410		TABLE 411		TABLE 412		TABLE 413		TABLE 414		TABLE 415		TABLE 416		TABLE 417		TABLE 418		TABLE 419		TABLE 420		TABLE 421		TABLE 422		TABLE 423		TABLE 424		TABLE 425		TABLE 426		TABLE 427		TABLE 428		TABLE 429		TABLE 430		TABLE 431		TABLE 432		TABLE 433		TABLE 434		TABLE 435		TABLE 436		TABLE 437		TABLE 438		TABLE 439		TABLE 440		TABLE 441	
---------------------------------	--	--------------------------------------	--	---------	--	---------	--	---------	--	---------	--	---------	--	---------	--	---------	--	---------	--	---------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--	-----------	--